Dilaton Quantum Gravity A Functional Renormalization Group Approach

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Outline

1 Introduction & Physical Context

- Nonrenormalizibility and Asymptotic Safety
- Challenges & Open Questions

2 FRG Analysis

- Flow Equations
- Fixed Point Solutions
- Scaling Solution



Renormalizibility and Quantum Gravity

• Einstein-Hilbert action

$$\Gamma[g_{\mu\nu}] = \frac{1}{16\pi G_N} \int d^d x \sqrt{g} \left(2\Lambda - R[g_{\mu\nu}]\right)$$

mass dimensions of the couplings

$$dim(\Lambda) = 2$$

 $dim(G_N) = 2 - c$

 \implies Perturbatively not renormalizable if d > 2.

The Flow Diagram of Quantum Einstein Gravity



Reuter, 1998

Coupling Constants approach a nontrivial UV Fixed Point

 \implies Prospect of Gravity being Asymptotically Safe

FRG Analysis

 $\underset{O}{\text{Summary and Outlook}}$

The Flow Diagram of Quantum Einstein Gravity



Christiansen, Litim, Pawlowski, Rodigast, 2012

Stable Infrared Scenarios

 \implies Prospect of UV and IR consistent theory

Asymptotically Safe Quantum Gravity

FRG Technicalities

- Regulator Dependence
- Background Dependence
- Truncation Stability

Coupling to SM

- Yang Mills Theory
- Background Dependence
- Asymptotic Freedom

Infrared Limit

- Trajectory $UV \rightarrow IR$
- GR as limiting case?

Hierarchy Problem

- $M_{SM} \approx 10^2 {
 m GeV}$
- $M_{Planck} pprox 10^{19} {
 m GeV}$

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Setup & Action

First investigated by Narain & Percacci, 2010 Extensions:

- investigation of IR limit
- trajectory UV \leftrightarrow IR
- improved UV analysis
- all investigations in d = 4 & deDonder gauge

Dilatation Symmetry

- Dilatations \leftrightarrow Conformal Transformations $g_{\mu\nu}(x) \mapsto \Omega(x)g_{\mu\nu}(x)$ with $\Omega = const$.
- \bullet Dilatation \leftrightarrow global resetting of the physical scale
- Dilatation Symmetry ↔ Physical Scale is introduced only by expectation value of the scalar field
- Arising Goldstone Boson: Dilaton

Dilatation Symmetric Actions

Physical Motivation



- a scale invariant IR limit for Einstein-Hilbert Quatum Gravity
- b generation of Planck mass via breaking of dilatation symmetry
- c quintessence cosmology scenarios with vanishing cosmological constant

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Fixed Point Action

Let
$$y = \frac{\chi^2}{k^2}$$
, $V_k(\chi^2) = k^4 y^2 v_k(y)$, $F_k(\chi^2) = k^2 y f_k(y)$

 \Rightarrow Dilatation Symmetric parts factored out

proposed fixed point for large y

$$\lim_{y \to \infty} f(y) = \xi \qquad \lim_{y \to \infty} v(y) = 0.$$

$$\Rightarrow \quad \Gamma = \int d^4 x \sqrt{g} \left(\frac{1}{2} g^{\mu\nu} \partial_{\mu} \chi \, \partial_{\nu} \chi - \frac{1}{2} \xi \chi^2 R \right) \qquad \text{(Jordan)}$$

$$\Rightarrow \quad \Gamma = \int d^4 x \sqrt{g} \left(\frac{1}{2} g^{\mu\nu} \partial_{\mu} \phi \, \partial_{\nu} \phi - \frac{1}{2} M^2 R \right) \qquad \text{(Einstein)}$$

Properties of the Fixed Point

• Let
$$f_0 = \lim_{y \to \infty} f_k(y)$$

- \Rightarrow Gravitational Interaction scales like $f_0^{-1}\chi^{-2}$
- \Rightarrow Gravity induced flow from $f_0^{-1}y^{-1}$ with $\lim_{y\to\infty} f_0^{-1}y^{-1} = 0$
- \Rightarrow For $\lim_{y\to\infty} v_k(y) = 0$ free scalar theory
 - $\bullet\,$ Flow of free scalar theory can at most produce a constant term $\propto\,v_{-2}$ in V and
 - For $g_N = 0$ no flow of gravitational interaction
- \Rightarrow Asymptotically

$$\lim_{y \to \infty} v(y) = v_{-2} y^{-2} + \dots$$
$$\lim_{y \to \infty} f(y) = \xi + f_{-1} y^{-1} + \dots$$

Flow Equation

- \$\mathcal{O}(R)\$ truncation
- no wave function renormalization
- curved background spacetime
- optimized cut-offs with $\Gamma_k^2(p^2) + \mathcal{R}_k(p^2) = \Gamma_k^2(p^2 + r_k(p^2))$ $r_k(p^2) = (k^2 - p^2) \times \Theta(k^2 - p^2)$
- set of flow equations at fixed y

$$\partial_t v_k(y) = 2 y v'_k(y) + \frac{1}{y^2} \zeta_V,$$

$$\partial_t f_k(y) = 2 y f'_k(y) + \frac{1}{y} \zeta_F.$$

• deDonder gauge fixing and mostly d = 4

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Flow Equation II

$$\begin{split} \zeta_{V} &= \frac{1}{192\pi^{2}} \left\{ 6 + \frac{30\,\tilde{V}}{\Sigma_{0}} + \frac{3(2\Sigma_{0} + 24\,y\,\tilde{F}'\,\Sigma_{0}' + \tilde{F}\Sigma_{1})}{\Delta} + \delta_{V}(\partial_{t}\tilde{F}) \right\} \\ \zeta_{F} &= \frac{1}{1152\pi^{2}} \left\{ 150 + \frac{30\,\tilde{F}\,(3\,\tilde{F} - 2\tilde{V})}{\Sigma_{0}^{2}} \\ &- \frac{12}{\Delta} \left(24\,y\,\tilde{F}'\,\Sigma_{0}' + 2\Sigma_{0} + \tilde{F}\Sigma_{1} \right) - 6y\,(3\,\tilde{F}'^{2} + 2\Sigma_{0}'^{2}) \\ &- \frac{36}{\Delta^{2}} \left[2y\,\Sigma_{0}\,\Sigma_{0}'\,(7\,\tilde{F}' - 2\tilde{V}')\,(\Sigma_{1} - 1) + 2\,\Sigma_{0}^{2}\,\Sigma_{2} \\ &+ 2\,y\Sigma_{1}\,(7\,\tilde{F}' - 2\tilde{V}')\,(2\,\Sigma_{0}\,\tilde{V}' - \tilde{V}\,\Sigma_{0}') \\ &+ 24\,y\,\tilde{F}'\,\Sigma_{0}\,\Sigma_{0}'\,\Sigma_{2} - 12\,y\,\tilde{F}\,\Sigma_{0}'^{2}\,\Sigma_{2} \right] + \delta_{F}(\partial_{t}\tilde{F}) \bigg\} \\ \tilde{V} &= y^{2}\,v_{k}(y) \;, \; \tilde{F} \; = \; y\,f_{k}(y), \end{split}$$

$$\begin{split} \Sigma_0 &= \frac{1}{2} \tilde{F} - \tilde{V} \ , \ \Delta &= \left(12 \, y \, \Sigma_0'^2 + \Sigma_0 \, \Sigma_1 \right), \\ \Sigma_1 &= 1 + 2 \, \tilde{V}' + 4 \, y \, \tilde{V}'' \ , \ \Sigma_2 &= \tilde{F}' + 2 \, y \, \tilde{F}'' \end{split}$$

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Fixed Point Solutions

Aim: Find Fixed Point Solutions

- a for y
 ightarrow 0 or ∞
 - Taylor Expansions in y or y^{-1}
 - free parameter $\xi = \lim_{y \to \infty} f_0$

Ultraviolet: $y \rightarrow 0$

- $\chi \to 0$ or $k \to \infty$
- Percacci at al, 2010

- Infrared: $y \to \infty$
 - $\chi \to \infty$ or $k \to 0$
 - current project, 2013

- b globally
 - connecting UV & IR
 - Padé ressumation
 - exponential improvement

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The Limit $y \to \infty$

• Expand

$$v(y) = v_0 + v_{-1}y^{-1} + v_{-2}y^{-2} + \dots,$$

 $f(y) = f_0 + f_{-1}y^{-1} + \dots$

- Finite limits for ζ_V and ζ_F
- \Rightarrow Flow of v_0 , v_{-1} and f_0 vanishes

• For
$$v_0 = v_{-1} = 0$$

 $\bar{\zeta}_V = \lim_{y \to \infty} \zeta_V = \frac{3}{32\pi^2} + \frac{5 + 33f_0}{96\pi^2(1 + 6f_0)} \frac{\partial_t f_0}{f_0}$
 $\bar{\zeta}_F = \lim_{y \to \infty} \zeta_F = \frac{77 + 534f_0}{192\pi^2(1 + 6f_0)} + \frac{17 + 186f_0 + 720f_0^2}{576\pi^2(1 + 6f_0)^2} \frac{\partial_t f_0}{f_0}$

- ${\ensuremath{\, \bullet }}$ closed set of flow equations guaranteed by derivative structure
- \Rightarrow fixed point solutions

$$\lim_{y\to\infty}f^*(y)=\xi+\frac{\bar{\zeta}_F}{2y}\ ,\ \ \lim_{y\to\infty}v^*(y)=\frac{\bar{\zeta}_V}{4y^2}.$$

The Limit $y \rightarrow 0$

Expand

$$\tilde{F} = yf = F_0 + F_1y + F_2y^2 + \dots$$

 $\tilde{V} = y^2v = V_0 + V_1y + V_2y^2 + \dots$

- system of flow equations not closed
- \Rightarrow Treat V_0 and F_0 as "free" parameters
 - extension from Narain, Percacci 2010
 - Apparent good convergence of the series

Closing the Gap: Possible Global Solutions

a Einstein-Hilbert Gravity + Scalar Field

$$\Leftrightarrow \tilde{V} = y^2 v(y) = 0.008620$$
 and $\tilde{F} = y f(y) = 0.04751$

b Continuations of Taylor Expansions around $y = \infty$



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Scaling Solution



• Taylor expansions around $y = \infty$ have $\epsilon \leq 10^{-2}$ for

 $y \leq y_0 \approx 1/(100\xi)$

• Brakedown of y^{-1} expansions expected for $y \rightarrow 0$

Improvement I: Padé ressumation

Strategy:

- stick with Taylor Solution for as long as possible
- smoothen solution for $y \rightarrow 0$
- connect to known solution for small y

Padé resummation:

- Expand both numerator and denominator in powers of y^{-1}
- dependency on parameter ξ remains



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Improvement II: Exponential Functions



Propagators Σ₀, Δ produce singularities in intermediate region
 See y as a coupling & use exponential parts ∝ exp(-c/y)

$$\tilde{F}_{exp}(y) = \tilde{F}_{Taylor}(y) + e^{-c/y}\tilde{F}_e(y); \quad \tilde{V}_{exp}(y) = \tilde{V}_{Taylor}(y) + e^{-c/y}\tilde{V}_e(y)$$

• parameter optimisation to connect solutions & minimise integrated error $\ensuremath{\mathcal{E}}$

Physical Implications

• Dilatation symmetric infrared limit due to $\lim_{V \to 0} V = 0, \quad \lim_{T \to 0} F = \xi \chi^2$

$$\lim_{y \to \infty} V = 0, \quad \lim_{y \to \infty} F = \xi_{y}^{2}$$

- Planck mass is generated by spontaneous breaking of Dilatation Symmetry via $\langle \chi \rangle = 0$
- \Rightarrow Realistic gravity possible on a pure fixed point trajectory
 - $\bullet\,$ quintessence cosmology for small deviations from fixed point and $y\gg 1$

$$\Gamma = \int d^4 x \sqrt{g} \left\{ -\frac{M^2}{2}R + \frac{1}{2}k^2(\varphi)\partial^{\mu}\varphi\partial_{\mu}\varphi + V(\varphi) \right\}$$
$$k^2(\varphi) = \frac{1}{4\xi} \left[1 + 6\xi + \left(\frac{M^2}{\bar{F}}\exp\left(\frac{\varphi}{M}\right) - 1\right)^{-1} \right]$$
$$V(\varphi) = \bar{V}\exp\left(-\frac{2\varphi}{M}\right)$$

 $\Rightarrow\,$ cosmological constant vanishes for $\varphi\rightarrow\infty$

Summary and Outlook

- prospect of scaling solution with scale invariant IR limit
- cosmological constant vanishes on and away from FP in the IR
- for small deviation from the fixed point trajectory admits quintessence cosmology

Thank you!