Spectral dimensions from the spectral action

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Outline

Motivation: Spontaneous Dimensional Reduction?

Generalised Spectral Dimensions

Spectral Action

Results

Conclusions



Is there spontaneous Dimensional Reduction in Short-Distance Quantum Gravity?

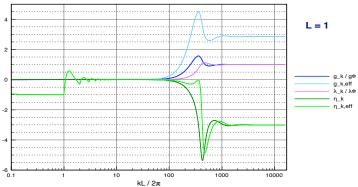
Important question in Quantum Gravity:

Structure of space-time at very short distances?

When lowering the distance scale / increasing the energy scale

- number of dimensions increases (as e.g. in the ADD model)?
- number of dimensions decreases (as e.g. in QEG or CDT)?
- volumes and areas become quantised (as e.g. in QLG)?
- space-time becomes discrete (as in the work presented here)?
- something completely else happens?

<u>Example:</u> Asymptotically safe (Quantum Einstein) gravity requires large anomalous dimensions at the UV FP!



RG flow in (G_N, Λ) truncation for four extended and one "large" dimension, *cf.* ADD model [N.A., Master Thesis 2014]

Quantum Einstein Gravity:

• In $D=4,5,6,\ldots$ the anomalous dimension related to the graviton wave function renormalisation becomes

$$\eta_N = 2 - D$$
.

[e.g. P. Fischer and D. Litim]

- This indicates a qualitative change of spacetime near and beyond Planck scale!
- Concept to describe this change: Spectral dimension D_S
- D_S = 2 close to UV FP
 [M. Reuter and F. Saueressig, JHEP 1112 (2011) 012.]

Fractal-like properties of a space(-time):

- Haussdorff dimension
 - Determined by # N of balls necessary to cover a set of points:

$$N(R) \propto 1/R^D$$

- Real line: $N(R) \propto 1/R$, i.e., D=1 coast of England: $D \approx 1.2$
- Spectral dimension D_s
 - Consider diffusion of scalar test particle on some manifold
 - ullet Average return probability ${\mathcal P}$ "feels" space-time dimension
 - $\mathcal{P} \propto T^{-D_S/2}$ with (fictitious) diffusion time T
- Spectral dimension = Hausdorff dim. of momentum space

Heuristic picture for the concept of a spectral dimension:

- Ping-pong ball sees three dimensions ©
- Table-football ball sees two dimensions ?@?
- ⇒ Interactions change the spectral dimension ©!

Spectral dimension near and beyond Planck scale:

Several ansätze for Quantum Gravity suggest $D_s = 2!$

Seen in

Causal Dynamical Triangulation & Asympt. Safe Gravity, arguments given for

Loop Quantum Gravity, High-Temperature Strings, Hořava-Lifshitz gravity & strong-coupling expansion of Wheeler-DeWitt equation.

[S. Carlip, talk in Petrópolis 2012 [arXiv:1207.4503] and arXiv:0909.3329

Proposed physical picture:

Focusing of geodesics (Asymptotic Silence)

Is the value $D_S=2$ in the UV generic for all approaches to Quantum Gravity?

Here:

Calculate the (classical) spectral dimension for Connes' non-commutative (resp., almost commutative) geometry at short distances.

Generalised Spectral Dimensions

Diffusion on classical manifold with metric g:

- Characterised by probability density P(x, x'; T)
- Average return probability $\mathcal{P}(T) = \frac{1}{V} \int d^d x \sqrt{g(x)} P(x, x' = x; T)$
- Spectral dimension $D_{\mathcal{S}}(T) = -2T \frac{\partial}{\partial T} \ln \mathcal{P}(T)$

Remarks

- Fictitious diffusion time T: Resolution scale $\mu \propto 1/\sqrt{T}$ [G. Calcagni, L. Modesto and G. Nardelli, arXiv:1408.0199 [hep-th]]
- Spectral dimension = Hausdorff dim. of momentum space

Generalised spectral dimensions

Calculation from 2-point function (= inverse propagator)

- (Generalised) Laplacian D^2 : $\partial_T P(x, x'; T) = D^2 P(x, x'; T)$
- For kinetic term $F(D^2)$: $\partial_T P(x, x'; T) = F(D^2) P(x, x'; T)$
- Average return probability $\mathcal{P}(T) = const.$ $\int_0^\infty dz \, z \, e^{-TF(z)}$
- Spectral dimension $D_S(T) = -2T \frac{\mathcal{P}'(T)}{\mathcal{P}(T)}$ with $\mathcal{P}'(T) = \partial_T \mathcal{P}(T)$

NB: If mass term exists then $D_S(T) = 2m^2T + D_S^{(0)}(T)$,

i.e. mass term contains no non-trivial information on propagation of particle.

 \implies Calculate $D_S^{(0)}(T)$ only. (Superscript omitted in the following.)

Spectral Action

Connes' non-commutative / almost commutative geometry:

[A. Connes, "Non-commutative geometry," Academic Press, 1994;

K. van den Dungen, W.D. van Suijlekom, Rev.Math.Phys.**24**, 1230004 (2012).]

- Spectral triple $\{A, \mathcal{H}, D\}$: Algebra, Hilbert space, Dirac operator
- Continuous spectral triple ⊗ discrete spectral triple
 (≡ Riemannian manifold) (non-comm. part → gauge sym.)
- 1 Generalisation of Riemannian geometry
- 2 Universal formula for the action of elementary fields
- (3) "Generating Functional" for Standard Model (SM) coupled to gravity at low scales (geometrical derivation of SM)
- 4 Very high energies: Framework for unification of SM & gravity

Spectral action

Spectral action for bosons:

$$S_{\chi,\Lambda} = \text{Tr}(\chi(D^2/\Lambda^2))$$

- \bullet χ positive function
- 2 D Dirac operator: contains spin 0, 1 & 2 fields
- 3 Λ physical scale, e.g., Planck mass or GUT Scale

Conjecture:

[M. A. Kurkov, F. Lizzi, D. Vassilevich, Phys. Lett. B 731 (2014) 311.]

"High energy bosons do not propagate"

... based on qualitative arguments ...

Spectral action

Dirac operator:

$$D^2 = -(\nabla^2 + E)$$

with endomorphism

$$E = -i\gamma^{\mu}\gamma_5\nabla_{\mu}\phi - \phi^2 - \frac{1}{4}R + \frac{i}{4}[\gamma^{\mu}, \gamma^{\nu}]F_{\mu\nu}.$$

- $oldsymbol{\phi}$ scalar field, $F_{\mu\nu}=\partial_{\mu}A_{\nu}-\partial_{\nu}A_{\mu}$ (Abelian) gauge field, R Ricci scalar
- 2 $\nabla_{\mu} = \nabla^{LC}_{\mu} + iA_{\mu}$ Levi-Cevita spin connection & gauge potential
- **3** Curvature $\Omega_{\mu\nu}:=[\nabla_{\mu},\nabla_{\nu}]=-\frac{1}{4}\gamma^{\rho}\gamma^{\sigma}R_{\rho\sigma\mu\nu}+iF_{\mu\nu}$



Expand Spectral Action for $\chi(z) = e^{-z}$ to 2nd order in fields:

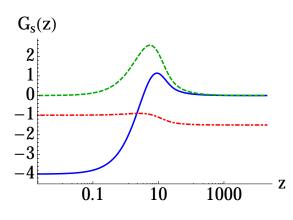
$$K^{(2)}(D^2, t) := \int d^4x \left[\phi F_0(-t\partial^2)\phi + F_{\mu\nu} \frac{1}{-2\partial^2} F_1(-t\partial^2) F_{\mu\nu} \right]$$
$$+ \Lambda^{-2} \mathbf{h}_{\mu\nu} F_2(-t\partial^2) \mathbf{h}_{\mu\nu} \quad \text{with} \quad t = 1/\Lambda^2$$

These functions are non-polynomial in $z = tp^2 = p^2/\Lambda^2$!

$$(4\pi)^{2}t F_{0}(z) =: G_{0}(z) = -1 + \frac{1}{2}zh(z)$$

$$(4\pi)^{2}t F_{1}(z) =: G_{1}(z) = -4 + 4h(z) + 2zh(z), \quad h(z) = \int_{0}^{1} d\alpha e^{-\alpha(1-\alpha)z}$$

$$(4\pi)^{2}t^{2}F_{2}(z) =: G_{2}(z) = -2 + h(z) + \frac{1}{4}zh(z)$$

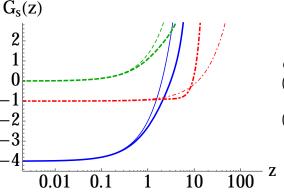


$$G_0(z = tp^2)$$
 scalar,
 $G_1(z)$ gauge field,
 $G_2(z)$ graviton.

$$\lim_{\substack{z\to\infty\\ \text{lim}\\ z\to\infty}} G_{0,1}=0$$

Effective Field Theory:

Expand in $t = 1/\Lambda^2$ to generate action polynomial in derivatives and **truncate** at some fixed order in $-t\partial^2$, resp., p^2/Λ^2 :



 $G_{\rm s}(z)$ at $\mathcal{O}(z^3)$ (thick lines) at $\mathcal{O}(z)$ (thin lines)



Effective Action for a generic function χ in $S_{\chi,\Lambda} = \text{Tr}(\chi(D^2/\Lambda^2))$, part quadratic in fields:

$$S_{\chi,\Lambda}^{(2)} = \frac{\Lambda^2}{(4\pi)^2} \int d^4x \left[\phi \, \mathcal{F}_{0,\chi} \left(-\partial^2/\Lambda^2 \right) \phi + A_\mu \, \mathcal{F}_{1,\chi} \left(-\partial^2/\Lambda^2 \right) A_\mu + \mathbf{h}_{\mu\nu} \, \mathcal{F}_{2,\chi} \left(-\partial^2/\Lambda^2 \right) \mathbf{h}_{\mu\nu} \right].$$

$$\mathcal{F}_{0,\chi}(z) = -Q_1 + \frac{Q_2}{2}z - \frac{1}{12}z + \frac{1}{120}z^2 + \dots, \qquad \text{s.t. } Q_n = 1 \text{ for }$$

$$\mathcal{F}_{1,\chi}(z) = \frac{4Q_0}{3}z - \frac{4Q_{-1}}{15}z^2 + \frac{Q_{-2}}{35}z^3 + \dots, \qquad \chi = e^z)$$

$$\mathcal{F}_{2,\chi}(z) = -Q_2 + \frac{Q_1}{12}z - \frac{Q_0}{40}z^2 + \frac{Q_{-1}}{336}z^3 + \dots \quad \star \quad Q_n'\text{s cannot be adjusted independently!}$$

 $-Q_1 + \frac{Q_0}{2}z - \frac{Q_{-1}}{12}z^2 + \frac{Q_{-2}}{120}z^3 + \dots$

 $\mathcal{F}_{0,\chi}(z) =$

 $\star \chi \to Q_n$ (unique,

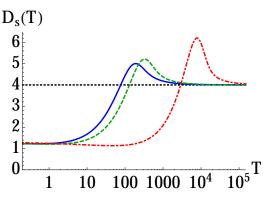
Possible truncations of the expansion:

Truncating the moments Q_n

- highest powers of $-\partial^2$ in matter and gravitational sector come with opposite signs
- ullet \Longrightarrow dynamical instability of the theory!

The effective field theory viewpoint

- \bullet truncate at a given power of the cutoff Λ
- valid for $-p^2/\Lambda^2 \ll 1$
- keep relevant and marginal operators
- \implies Q_n 's can be adjusted s.t. all propagators are stable!



F₀ scalar,F₁ gauge field,F₂ graviton.

$$\frac{\mathsf{IR}\;(T\to\infty):}{Dc\to 4}$$

$$\frac{{
m UV}~(T
ightarrow 0):}{D_S(0)=4/N_{max}}$$
 determined by highest power of truncation!

$$D_S(0) \rightarrow 0$$
 for $N_{max} \rightarrow \infty$

Include full momentum dependence of inverse propagators:

• Regularise occurring momentum integrals

$$P(T; \Lambda_{\mathrm{UV}}) = \int^{\Lambda_{\mathrm{UV}}} \frac{d^4p}{(2\pi)^4} e^{-TF(p^2)}$$

- $\Lambda_{\mathrm{UV}} \to \infty$ (NB: Λ fixed!)
- Based on the late-time expansion of the occurring functions one can analytically show

$$D_S(T) = \lim_{\Lambda_{\mathrm{UV}} \to \infty} D_S(T; \Lambda_{\mathrm{UV}}) = 0.$$

Generic function χ in spectral action: Non-polynomial in derivatives!

Consequence:

Spectral dimensions vanish in UV / for very small distances! Space-time fractures into non-communicating points!

Conclusion

- Calculation of the scale-dependent spectral dimensions of bosons from a spectral action.
- Spectral dimensions vanish!
 ("High-energy bosons do not propagate!")
- Space-time fractures into non-communicating points!
- ? Different than in all other approaches to Quantum Gravity!
- ? Vanishing spectral dimension for all non-polynomial actions?
- ? Conditions on function χ defining bosonic spectral action?
- ? Spectral dimensions if quantum propagators are used?