# Local Quantum Gravity

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IMPRS *ptfs* 

#### Introduction

- Motivation
- Truncation

#### 2 Momentum locality

- Definition and examples
- Proof for graviton two- and three-point function

#### 3 Fixed points and phase diagram

- Investigate the asymptotic safety scenario with a fully dynamical, minimally self-consistent ansatz
- Newton's coupling from graviton three-point function
- No background-field approximation

$$\Gamma_k[\overline{g},\phi] \neq \Gamma_k[\overline{g}+\phi] \tag{1}$$

• Vertex expansion to disentangle background and dynamical fields

$$\Gamma_{k}[\bar{g},\phi] = \sum_{n} \frac{1}{n!} \Gamma^{(\phi_{1}\dots\phi_{n})}[\bar{g},0] \phi_{1}\dots\phi_{n} \quad \text{with} \quad \phi = (h,c,\overline{c})$$

Graviton two-point function:

$$\partial_t (-)^{-1} = -\frac{1}{2} + -2 + -2$$

Graviton three-point function:



RG-invariant ansatz for *n*-point function:

$$\Gamma^{(\phi_1...\phi_n)} = \prod_{i=1}^n \sqrt{Z_{\phi_i}(p_i^2)} G_n^{\frac{n}{2}-1} \mathcal{T}^{(n)}(p_1,\ldots,p_n;\Lambda_n)$$
(2)

 $\mathcal{T}^{(n)}$  contains classical Einstein-Hilbert tensor structures

Graviton two-point function:

$$\Gamma^{(hh)} = Z_h(p^2) \mathcal{T}^{(2)} \sim Z_h(p^2)(p^2 - 2\Lambda_2)\Pi_{ au au} + ext{other tensor structures}$$

Graviton three-point function:

$$\Gamma^{(hhh)} = \sqrt{Z_h(p_1^2)Z_h(p_2^2)Z_h(p_3^2)}G_3^{1/2}\mathcal{T}^{(3)}(p_1, p_2, p_3; \Lambda_3)$$

RG-invariant ansatz for *n*-point function:

$$\Gamma^{(\phi_1...\phi_n)} = \prod_{i=1}^n \sqrt{Z_{\phi_i}(p_i^2)} G_n^{\frac{n}{2}-1} \mathcal{T}^{(n)}(p_1,\ldots,p_n;\Lambda_n)$$
(2)

 $\mathcal{T}^{(n)}$  contains classical Einstein-Hilbert tensor structures

Graviton two-point function:

 $\Gamma^{(hh)} = Z_h(p^2) \mathcal{T}^{(2)} \sim Z_h(p^2)(p^2 + M^2) \Pi_{TT}$  + other tensor structures

 $M^2 = -2\Lambda_2 = \text{graviton mass parameter}$ 

We focus on the transverse-traceless part of the tensor structure:

• Split up into momentum dependent and momentum independent part

$$\mathcal{T}^{(n)}(p_1,\ldots,p_n;\Lambda_n) = \Lambda_n \mathcal{T}^{(n)}(0;1) + \mathcal{T}^{(n)}(p_1,\ldots,p_n;0)$$
 (3)

- Contract with *n* transverse-traceless projectors
- Obtain two operators to project on flow of  $\Lambda_n$  and  $G_n$
- RHS of flow equation

$$\mathsf{Flow}_{\Lambda}^{(3)} := \frac{\dot{\Gamma}^{(3)}}{Z_{h}^{3/2}(\rho^{2})} \circ \Pi_{\mathsf{TT}}^{3} \circ \mathcal{T}^{(3)}(0;1) \tag{4}$$

Graviton three-point function depends on  $p_1$ ,  $p_2$  and  $p_3$ 

- Momentum conservation: 3 degrees of freedom left  $(|p_1|, |p_2| \text{ and } \vartheta_{12})$
- Use maximally symmetric momentum configuration

$$p := |p_1| = |p_2|$$
  $\vartheta_{12} = 2\pi/3$  (5)

• Only dependence on scalar p is left

# Momentum projection

• Bilocal projection for  $G_n$ 

$$\dot{g}_{3} = (2 + 3\eta_{h}(k^{2}))g_{3} - \frac{24}{19}(\eta_{h}(k^{2}) - \eta_{h}(0))\lambda_{3}g_{3} + 2N_{g}\sqrt{g_{3}}k\left(\mathsf{Flow}_{G}^{(3)}(k^{2}) - \mathsf{Flow}_{G}^{(3)}(0)\right)$$
(6)

• 
$$p = 0$$
 projection for  $\Lambda_n$ 

$$\dot{\lambda}_3 = \left(\frac{3}{2}\eta_h(0) - 1 - \frac{\dot{g}_3}{2g_3}\right)\lambda_3 + \frac{\mathcal{N}_\lambda}{\sqrt{g_3}k}\mathsf{Flow}^{(3)}_\Lambda(0) \tag{7}$$

• Full momentum dependence of  $\eta_h(p^2)$  and  $\eta_c(p^2)$  (Christiansen et al. 2014)

# Truncation

#### • Flat Euclidean background

$$\bar{g}_{\mu\nu} = \delta_{\mu\nu} + h_{\mu\nu}$$

• Closure of flow equations

$$G_n(p_1,\ldots,p_n)\equiv G_3=:G$$
  
 $\Lambda_5\equiv\Lambda_4\equiv\Lambda_3$ 

• Parameters of the truncation

$$G, \Lambda_3, M^2, \eta_h(p^2), \eta_c(p^2)$$

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Fixed points and phase diagram

# Questions so far?

# Investigation of momentum dependence



Unexpected momentum dependence For graviton two-point function see *Christiansen et al. 2012* 

# Definition of momentum locality for an *n*-point vertex

$$\lim_{t_i/k^2 \to \infty} \frac{|\dot{\Gamma}_k^{(n)}(p_1,...,p_n)|}{|\Gamma_k^{(n)}(p_1,...,p_n)|} = 0$$

 $t_i$  = momentum channels

(each propagator must carry  $p \to \infty$ )

|.| = projection on tensor structure

- Implements the separation of scales
- An RG-step in the IR does not change physics in the UV

# Examples

Perturbatively renormalisable theories

Example:  $S = \int_x \frac{1}{2}m^2\phi^2 + \frac{1}{2}(\partial\phi)^2 + \lambda_1\phi^4$ 



#### Momentum local!

Applies to standard theories like gauge-theories in four dimensions

# Examples

Perturbatively non-renormalisable theories

Example: 
$$S = \int_x \frac{1}{2}m^2\phi^2 + \frac{1}{2}(\partial\phi)^2 + \lambda_1\phi^4 + \lambda_2\phi^2(\partial\phi)^2$$



#### Not momentum local!

### Graviton two-point function

Truncation: 
$$S \sim \int_{X} \sqrt{g} (2\Lambda - R) + \dots$$
  
 $\Gamma^{(hh)} \sim p^2$ 
 $\Gamma^{(hhh)} \sim p^2$ 
 $\Gamma^{(hhh)} \sim p^2$ 
 $\Gamma^{(5h)} \sim p^2$ 
 $\dot{\Gamma}^{(5h)} \sim p^2$ 

#### Non-trivially momentum local!

## Graviton three-point function

Truncation: 
$$S \sim \int_x \sqrt{g} (2\Lambda - R) + \dots$$
  
 $\Gamma^{(hh)} \sim p^2 \qquad \Gamma^{(hhh)} \sim p^2 \qquad \Gamma^{(4h)} \sim p^2 \qquad \Gamma^{(5h)} \sim p^2$ 



$$\lim_{p_i/k\to\infty}\frac{\dot{\Gamma}^{(hhh)}}{\Gamma^{(hhh)}}=0$$

#### Non-trivially momentum local!

Holds for all momentum configurations with  $p_i 
ightarrow \infty$ 

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#### Graviton three-point function



Link to diffeomorphism invariance!

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# Phase diagram



 $F_1$ : UV fixed point  $F_2$ : Gaußian fixed point  $F_3$ : IR fixed point



(Analytic approximation)

UV fixed point properties:

- $(g_{\text{UV}}^*, \mu_{\text{UV}}^*, \lambda_{3,\text{UV}}^*) = (0.66, -0.59, 0.11)$
- $(\theta_{1/2}, \theta_3) = (-1.4 \pm 4.1 i, 14)$
- Two attractive and one repulsive direction
- Two dimensional critical hypersurface
- Repulsive direction approximately along  $\lambda_3$ -axis



(Analytic approximation)

IR fixed point properties:

- $(g_{\mathsf{IR}}^*, \mu_{\mathsf{IR}}^*, \lambda_{\mathsf{3},\mathsf{IR}}^*) = (0, \infty, -\infty)$
- Classical scaling close to IR fixed point
- Non-trivial IR fixed point at  $(0, -1, -\infty)$  not found (Christiansen et al. 2014)
- Expansion about non-flat background required

Current approximation:

$$G_n(p_1,\ldots,p_n)\equiv G_3$$

Is that a good approximation?

(I.e. does  $Z_h(p^2)$  capture higher order momentum dependence?)

Compare full momentum dependence of LHS with RHS (at the fixed point)

LHS = 
$$\partial_t \left( Z_h^{3/2}(p^2) G^{1/2} \left( \frac{171}{32} p^2 + \frac{9}{4} \Lambda_3 \right) \right) \stackrel{?}{=} \operatorname{Flow}_G^{(3)}(p^2) = \operatorname{RHS}$$
 (8)

# Higher order momentum dependence



Setup:

- Genuine Newton's coupling from graviton three-point function
- Minimally self-consistent system

Results:

- New notion of momentum locality
- Proof of momentum locality for graviton 2- and 3-point function
- UV fixed point, which supports the asymptotic safety scenario
- IR fixed point with classical scaling

- Inclusion of matter degrees of freedom (see 'Asymptotic safety of gravity-matter systems' J. Meibohm, J. M. Pawlowski and MR arXiv:1510.07018)
- Graviton four-point function
- Expansion around non-flat background