

A link that matters:
Towards phenomenological constraints
on quantum gravity models

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In collaboration with: A. Eichhorn and A.D. Pereira

Motivation

Towards Quantum Gravity Phenomenology

How to confront quantum gravity with experiments?

- Direct probes seems to be far away from the current technology
- Indirect probes might contain some imprints from quantum gravity
 - Cosmological observations?
 - Astrophysical objects?
 - Particle physics experiments?

ASQG and Matter Interactions

A link that matters: From Planck to TeV

Basic idea: Use the RG flow to connect UV fixed points with IR physics

- UV completion for SM-like couplings?
- Predicting values at TeV scale?

ASQG and Matter Interactions

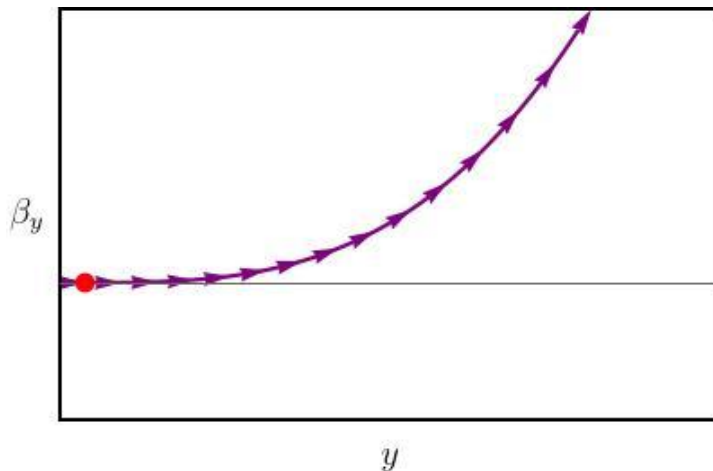
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Exemple: Yukawa Interaction

$$\beta_y = \# y^3$$



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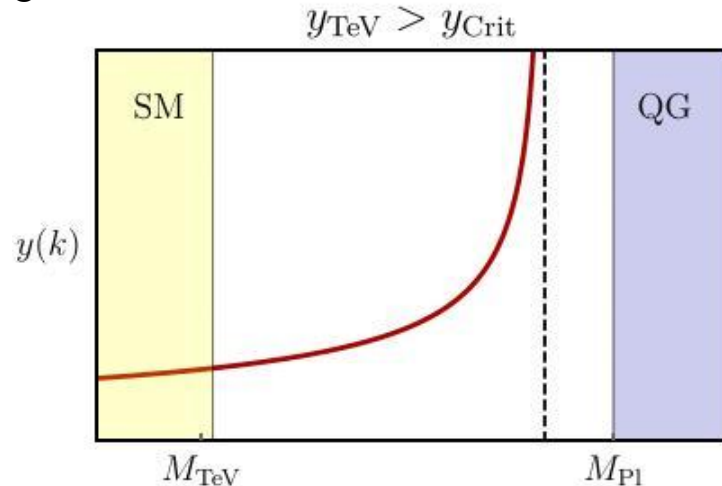
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- Landau pole!
- New physics below Planck scale?



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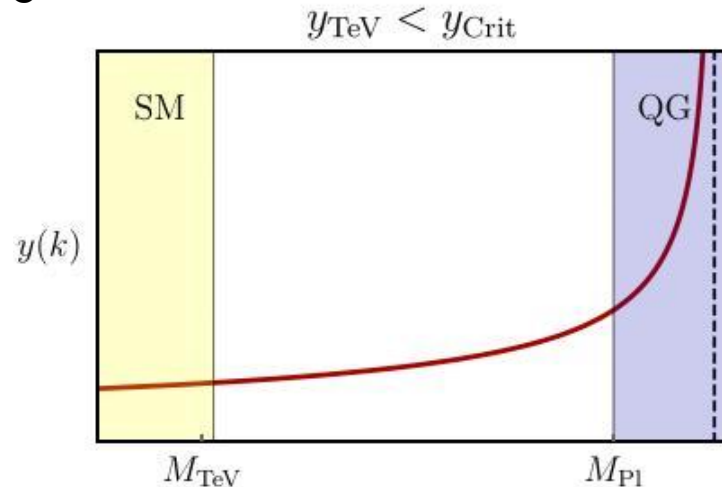
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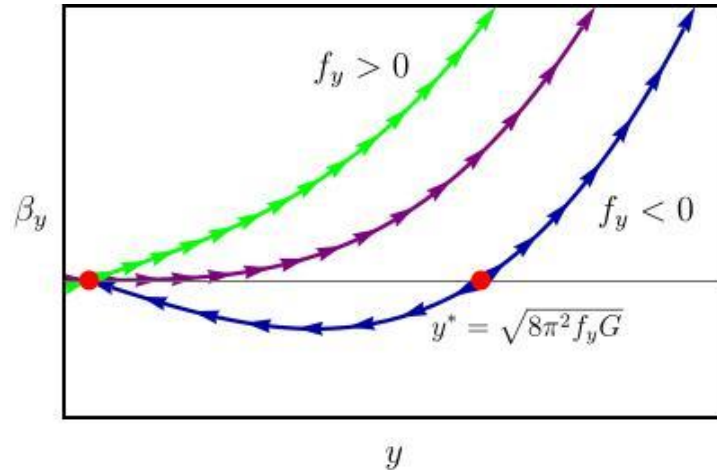
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Exemple: Yukawa Interaction

$$\beta_y = \# y^3 + f_y G y$$

- $f_y > 0$ - screening
- $f_y < 0$ - anti-screening contribution



ASQG and Matter Interactions

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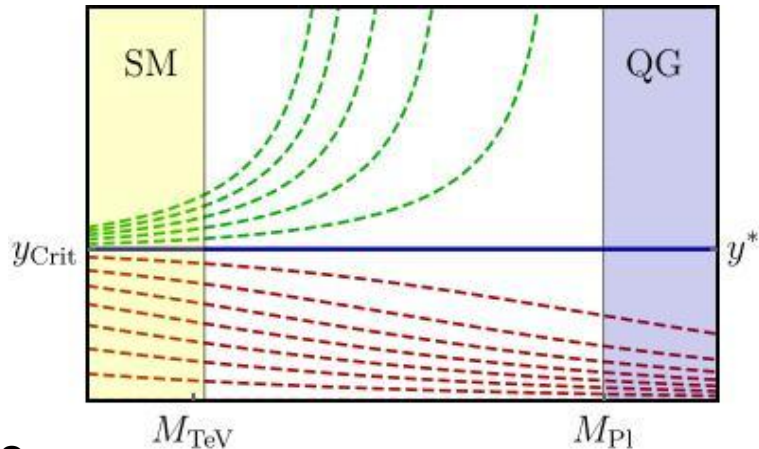
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Exemple: Yukawa Interaction

Yukawa flow with $f_y < 0$

- $y_{\text{TeV}} < y_{\text{Crit}}$ - A.F. trajectories
- $y_{\text{TeV}} = y_{\text{Crit}}$ - Safe (predictive) trajectories
- $y_{\text{TeV}} > y_{\text{Crit}}$ - Unsafe trajectories



ASQG and Matter Interactions

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Developments in ASQG + SM

- Higgs mass predictions Shaposhnikov and Wetterich, '09
- UV completion in the hypercharge sector Christiansen and Eichhorn '17
Eichhorn and Versteegen '17
- Top-Bottom mass difference Eichhorn and Held '17, '18, '19

ASQG and Matter Interactions

Constraining different candidates for ASQG

Consistency tests for Quantum Gravity candidates:

- Resolving the Landau pole problem for SM-like couplings?
- Viable “predictions” for SM-like couplings at TeV?

ASQG and Matter Interactions

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Metric based candidates for ASQG

- “Standard” ASQG (Diff. Symmetry)
- Quantum Unimodular Gravity (TDiff. Symmetry)
- Weyl squared gravity (Weyl + Diff. Symmetry)

Unimodular (Quantum) Gravity

Basics of Unimodular Gravity

GR with fixed metric determinant: $\det g_{\mu\nu} = \omega$

$$S_{\text{UG}}[g_{\mu\nu}] = -\frac{1}{16\pi G_N} \int d^4x \sqrt{\omega} R(g) \implies R_{\mu\nu} - \frac{1}{4}g_{\mu\nu}R = -8\pi G_N \left(T_{\mu\nu} - \frac{1}{4}g_{\mu\nu}T \right)$$

- Classically equivalent to GR
- C.C. arises as an integration constant
 - Different perspective on the C.C. problem
- UG is symmetric under TDiff. rather than Diff.
 - Symmetry group in the bottom-up approach

Weinberg '89
Percacci '18

Van Der Bij, Van Dam and
Yee Jack Ng '82

Unimodular (Quantum) Gravity

Quantum Unimodular Gravity

Quantizing Unimodular Gravity:

$$S_{\text{UG}}[g_{\mu\nu}] = -\frac{1}{16\pi G_N} \int d^4x \sqrt{\omega} R(g) \quad \Rightarrow \quad Z_{\text{UG}} \sim \int_{\text{UG}} \mathcal{D}g_{\mu\nu} e^{iS_{\text{UG}}[g_{\mu\nu}]}$$

- Quantum (in)equivalence with GR?
- How to implement unimodularity in the path integral?
 - Field redefinition?
 - Lagrange multipliers?
- Our approach:

$$g_{\mu\nu} = \bar{g}_{\mu\alpha} (e^h)^\alpha{}_\nu \quad \text{with unimodular background and traceless fluctuations}$$

Gravity-Matter Systems in Unimodular Setting

Truncation:

$$\Gamma_k = \Gamma_k^{\text{UG}} + \Gamma_k^{\text{SM-like}}$$

→ (Unimodular) Gravitational Sector

$$\Gamma_k^{\text{UG}} = \frac{1}{16\pi G_N} \int_x \sqrt{\omega} (-R + \bar{a}R^2 + \bar{b}R_{\mu\nu}^2) + \Gamma_k^{\text{g.f.}}$$

→ SM-like sector

$$\begin{aligned} \Gamma_k^{\text{SM-like}} &= \frac{Z_A}{4} \int_x \sqrt{\omega} g^{\mu\alpha} g^{\nu\beta} F_{\mu\nu} F_{\alpha\beta} + \Gamma_k^{\text{g.f.}} \\ &+ \frac{1}{2} \int_x \sqrt{\omega} (Z_{k,\phi} g^{\mu\nu} \partial_\mu \phi \partial_\nu \phi + 2V_k(\phi)) + \int_x \sqrt{\omega} (Z_{k,\psi} i\bar{\psi} \gamma \cdot D \psi + iy_k \phi \bar{\psi} \psi) \end{aligned}$$

Gravity-Matter Systems in Unimodular Setting

Yukawa and abelian gauge couplings:

$$\beta_y|_{\text{grav}} = f_y G y \quad \Rightarrow \quad f_y = \frac{1}{160\pi} \left(\frac{75(2+3b)}{(1+b)^2} + \frac{2(9-42a-14b)}{(1-6a-2b)^2} \right)$$

$$\beta_{g_Y^2}|_{\text{grav}} = f_{g_Y^2} G g_Y^2 \quad \Rightarrow \quad f_{g_Y^2} = -\frac{1}{90\pi} \left(\frac{5(10+7b)}{(1+b)^2} - \frac{4(5-21a-7b)}{(1-6a-2b)^2} \right)$$

Gravity-Matter Systems in Unimodular Setting

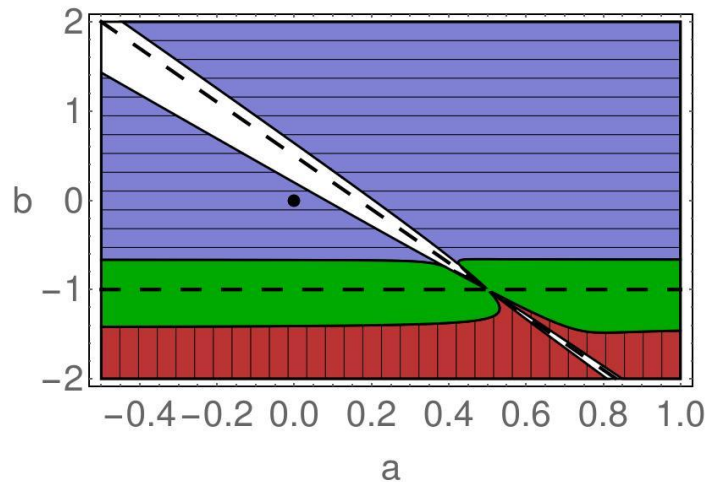
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Quantum gravity anti-screening contributions:

- $f_y < 0$ and $f_{g_Y^2} < 0$
- Regions for UV completion and predictivity are available!



Gravity-Matter Systems in Unimodular Setting

Higgs mass predictivity:

In UG, the gravitational contribution to the scalar potential flow comes exclusively from the scalar anomalous dimension

$$V_k(\phi) = \sum_n k^{4-2n} \lambda_n \phi^{2n} \implies \beta_{\lambda_n}|_{\text{grav}} = n \eta_{\phi}|_{\text{grav}} \lambda_n$$

$$\eta_{\phi}|_{\text{grav}} = \frac{1}{40\pi} \left(\frac{25(2+3b)}{(1+b)^2} + \frac{4(5-33a-11b)}{(1-6a-2b)^2} \right)$$

Gravity-Matter Systems in Unimodular Setting

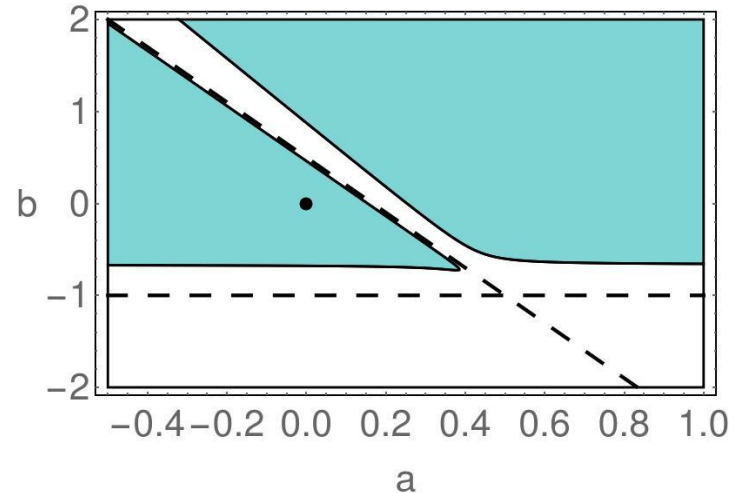
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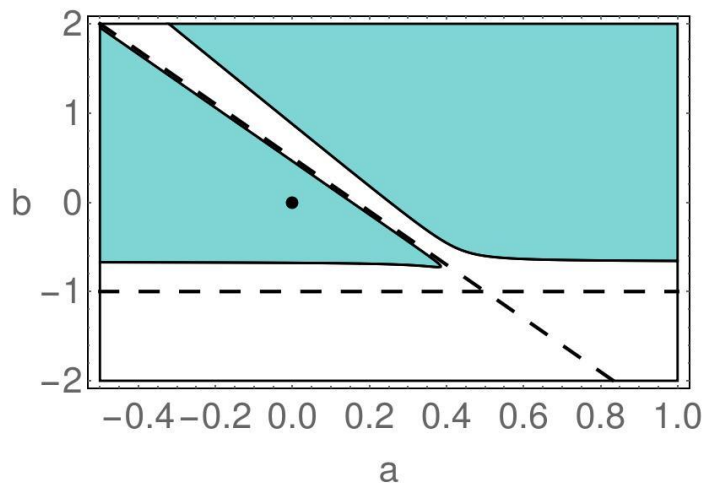
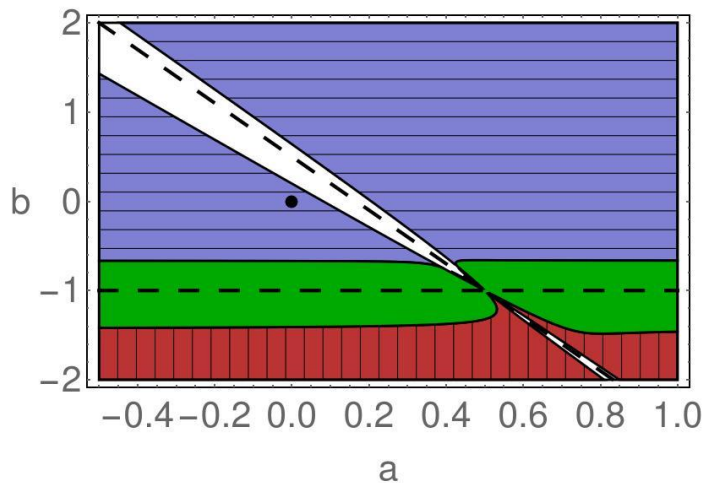
Predictivity of the Higgs mass requires positive contribution to the anomalous dimension



Gravity-Matter Systems in Unimodular Setting

Combined SM-like couplings:

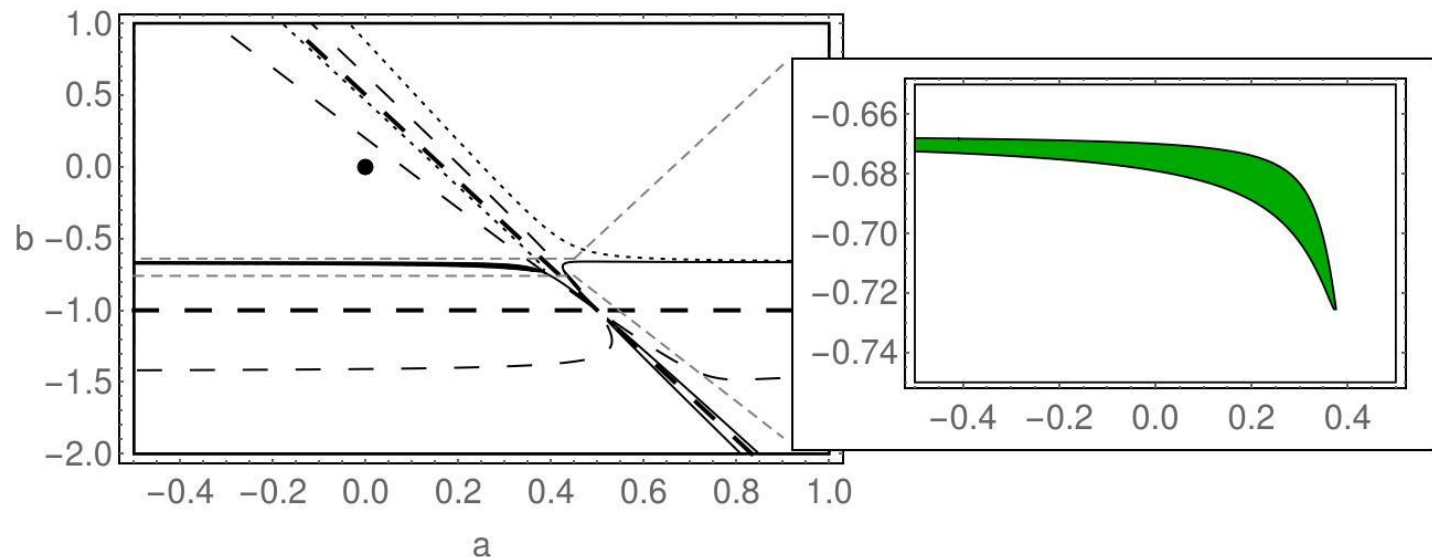
Simultaneous UV completion and Higgs mass predictivity?



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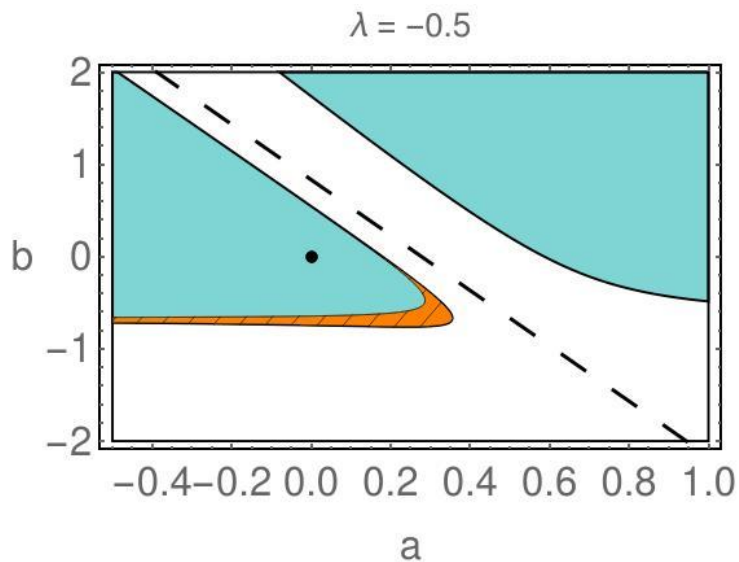
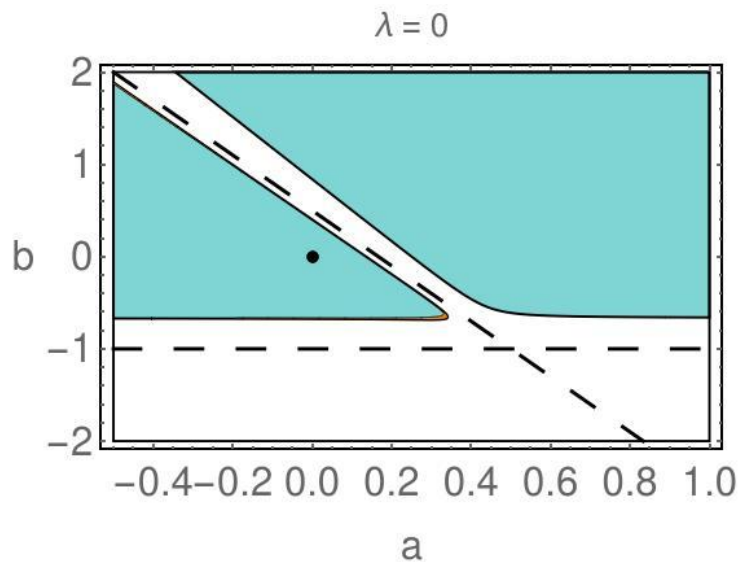
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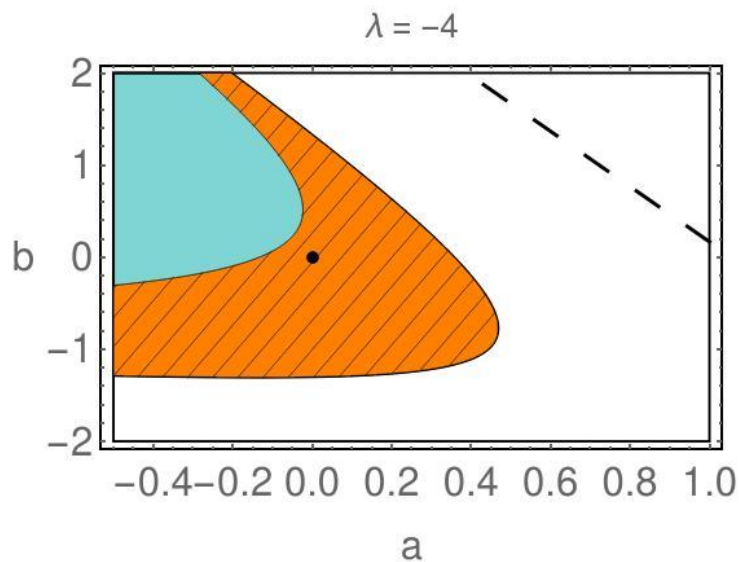
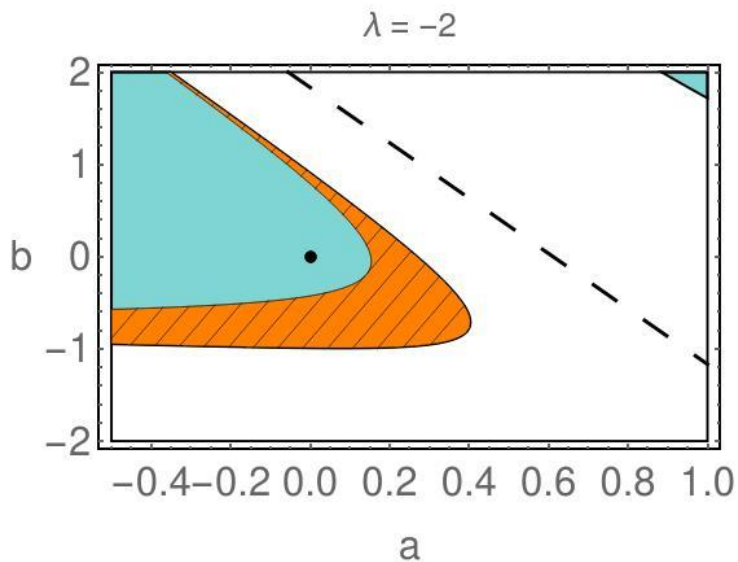
Comparison with “Standard” ASQG

UV completion and Higgs mass predictivity?



Comparison with “Standard” ASQG

UV completion and Higgs mass predictivity?



→ The overlap viable region for UV completion/predictivity of SM-like couplings is enlarged by negative values of C.C.

Weyl-Squared Gravity

Exploring larger symmetries in Quantum Gravity

Weyl-squared gravity is characterized by an enhanced symmetry corresponding to Weyl transformations:

$$g_{\mu\nu}(x) \mapsto g'_{\mu\nu}(x) = \Omega^2(x) g_{\mu\nu}(x)$$

For a review see: [Scholz '18](#)

→ Strong restrictions on the allowed (local) interactions:

$$S_{\text{Weyl}} = \frac{1}{2w} \int_x \sqrt{g} C_{\mu\nu\alpha\beta}^2$$

→ Power-counting renormalizable and asymptotically-free

- Natural candidate for UV completion [Fradkin and Tseytlin '82](#)

→ Issues related with unitarity to be understood

Weyl-Squared Gravity

The interplay gravity-matter in Weyl-Squared setting

- UV completion/predictivity of SM-like couplings?
- Universality of quantum-gravity contributions due to dimensionless couplings in the gravitational sector?

FRG truncation

$$\Gamma_k^{\text{WG}} = \int_x \sqrt{g} \left(\frac{1}{2w} C_{\mu\nu\alpha\beta}^2 + \frac{Z_{k,\phi}}{2} g^{\mu\nu} \partial_\mu \phi \partial_\nu \phi + \frac{\lambda_{k,4}}{4!} \phi^4 + \frac{\chi}{6} \phi^2 R \right. \\ \left. + \frac{Z_{k,A}}{4} g^{\mu\alpha} g^{\nu\beta} F_{\mu\nu} F_{\alpha\beta} + Z_{k,\psi} i \bar{\psi} \gamma \cdot D \psi + i y_k \phi \bar{\psi} \psi \right) + \Gamma_k^{\text{g.f.}}$$

Weyl-Squared Gravity

UV completion/predictivity of SM-like couplings?

Yukawa and abelian gauge sectors:

$$\beta_y|_{\text{WG}} = f_y w y \quad \Rightarrow \quad f_y = \frac{15}{64\pi^2} \Phi_2^3$$

$$\beta_{e_Y^2}|_{\text{WG}} = f_{e_Y^2} w e_Y^2 \quad \Rightarrow \quad f_{e_Y^2} = \frac{5}{12\pi^2} (\Phi_2^3 - 3\Phi_3^4)$$

Threshold integrals: $\Phi_n^p = \frac{1}{\Gamma(n)} \int_0^\infty dy y^{n-1} \frac{r(r) - yr'(y)}{(y + r(y))^p}$

Weyl-Squared Gravity

UV completion/predictivity of SM-like couplings?

Yukawa and abelian gauge sectors:

$$\beta_y|_{\text{WG}} = f_y w y \quad \Rightarrow \quad f_y = \frac{15}{64\pi^2} \Phi_2^3 = \frac{15}{128\pi^2}$$

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Threshold integrals: $\Phi_n^p = \frac{1}{\Gamma(n)} \int_0^\infty dy y^{n-1} \frac{r(r) - yr'(y)}{(y + r(y))^p}$

→ Universal result: $\Phi_n^{n+1} = \frac{1}{\Gamma(n+1)}$

→ The results are also universal with respect to field parametrization

Ohta and Percacci '16

Weyl-Squared Gravity

UV completion/predictivity of SM-like couplings?

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UV completion/predictivity?

- UV completion in the Yukawa sector depend on the sign of w
- No UV completion induced in the abelian gauge sector

Concluding Remarks

- The interplay between gravity and matter might help to pave some way to perform phenomenological tests in quantum gravity
 - UV completion/predictivity of SM-like couplings as viability tests
- Unimodular gravity-matter systems
 - Quite similar results in comparison with standard ASQG with vanishing C.C.
 - Quantum UG seems to be less predictive than standard ASQG with C.C.
 - Possible to mimic in UG with the introductions of “graviton mass parameters” motivated by mSTI’s
 - Further analysis is necessary in order to get quantitative “predictions” in UG
- Weyl-squared gravity
 - No UV completion in the abelian gauge sector
 - Universal results with respect to the choice of regulator and field parametrization

Acknowledgements

- CNPq and Capes for the financial support
- ITP - Heidelberg University for the hospitality

Thank you for your attention!